

First order ordinary differential equations

↑
only first
derivatives

↑
only one independent variable
(usually x or t)

① Exponential decay

Radioactivity = decay of atomic nuclei

Key experimental fact: the probability that a nucleus will decay is independent of how long it has existed

Let $N(t)$ = # of radioactive nuclei present in a sample at time t

Define decay constant λ as probability per unit time that any one nucleus will decay
(called constant because indep of time)

Define activity $A(t)$ as number of nuclei in sample that decay per unit time [what you actually measure]

$$A(t) = \lambda N(t)$$

This reduces the number of nuclei present

$$\frac{dN}{dt} = -A = -\lambda N$$

$$\boxed{\frac{dN}{dt} + \lambda N = 0} \quad \text{decay equation}$$

A separable 1st order o.d.e [is one that] can be rearranged so [that] all occurrences of the dependent variable appear on one side of the eqn, and all occurrences of the independent variable appear on the other.

It can be solved by integrating both sides.

$$\frac{dN}{dt} = -\lambda N$$

$$\frac{dN}{N} = -\lambda dt$$

const of integration

$$\ln N = -\lambda t + C_1$$

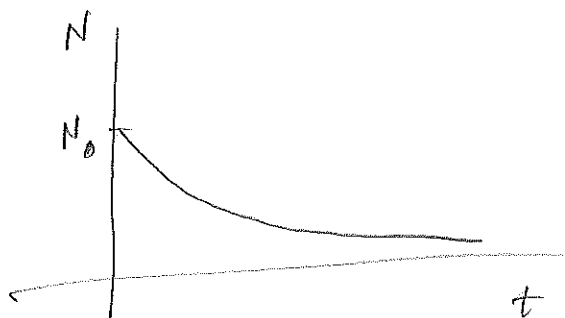
The solution of a 1st order eqn always contains one constant undetermined by the eqn. It can be determined using initial (or boundary) conditions.

Let $N_0 = \# \text{ nuclei at } t=0 \Rightarrow C_1 = \ln N_0$

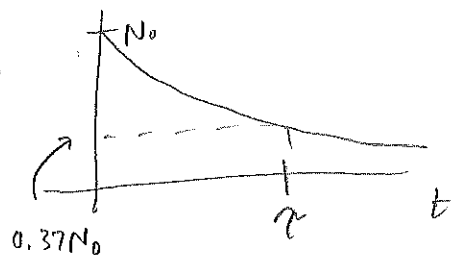
$$\ln N - \ln N_0 = -\lambda t$$

$$\ln\left(\frac{N}{N_0}\right)$$

$$N(t) = N_0 e^{-\lambda t}$$



Define $\tau = \frac{1}{\lambda} \Rightarrow N(t) = N_0 e^{-\frac{t}{\tau}}$



$$N(\tau) = N_0 e^{-1} \approx 0.37 N_0$$

so after time τ only 37% of nuclei remain

(one can show τ = mean life of a nucleus)

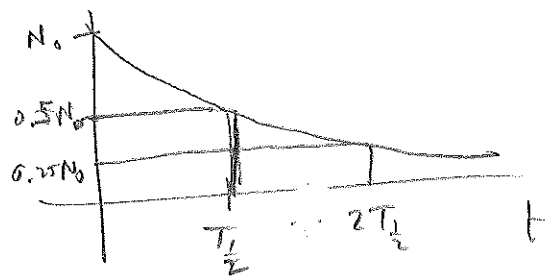
→ see how to plot in next lecture

Define half-life $T_{1/2}$ by $N(T_{1/2}) = \frac{1}{2} N_0$

(time it takes half of nuclei to decay)

Since $N(T_{1/2}) = N_0 e^{-T_{1/2}/\tau}$

$$\Rightarrow T_{1/2} = \tau \ln 2 \approx 0.69 \tau$$

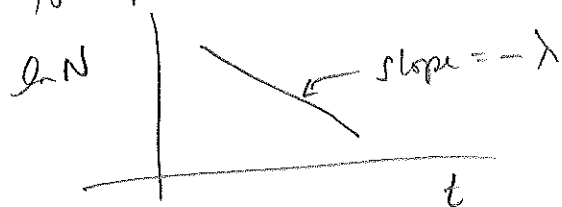


Prove to yourself that

$$N = N_0 \cdot \left(\frac{1}{2}\right)^{t/T_{1/2}}$$

so that for each half-life, only half of nuclei remain

To measure τ , do a semilog plot of data

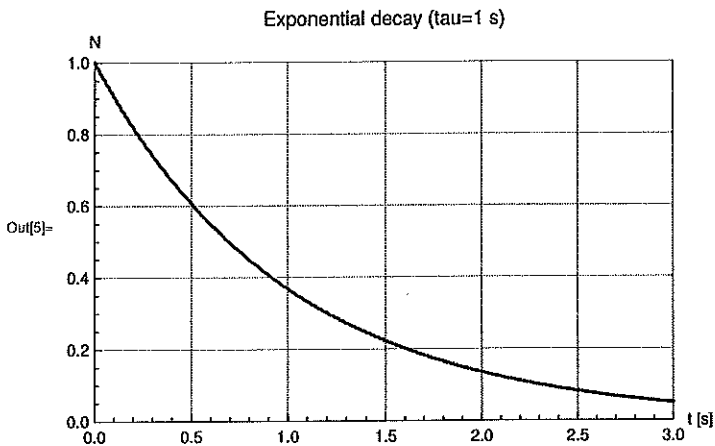


$$\ln N = \ln N_0 - \lambda t$$

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In[5]:= Plot[ Exp[-x], {x, 0, 3},
  PlotRange -> { {0, 3}, {0, 1.0} },
  GridLines -> Automatic,
  AxesLabel -> {"t [s]", "N"},
  PlotLabel -> "Exponential decay (tau=1 s)"]

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In[6]:= Export["ExpDecay.pdf", %]

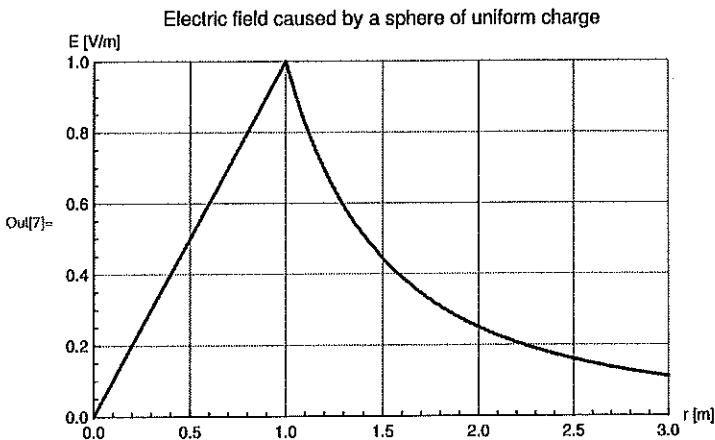
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Out[6]= ExpDecay.pdf

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In[7]:= Plot[Piecewise[{{r, r < 1}, {1/r^2, r > 1}}, {r, 0, 3}, PlotRange -> { {0, 3}, {0, 1} },
  GridLines -> Automatic,
  AxesLabel -> {"r [m]", "E [V/m]"},
  PlotLabel -> "Electric field caused by a sphere of uniform charge"]

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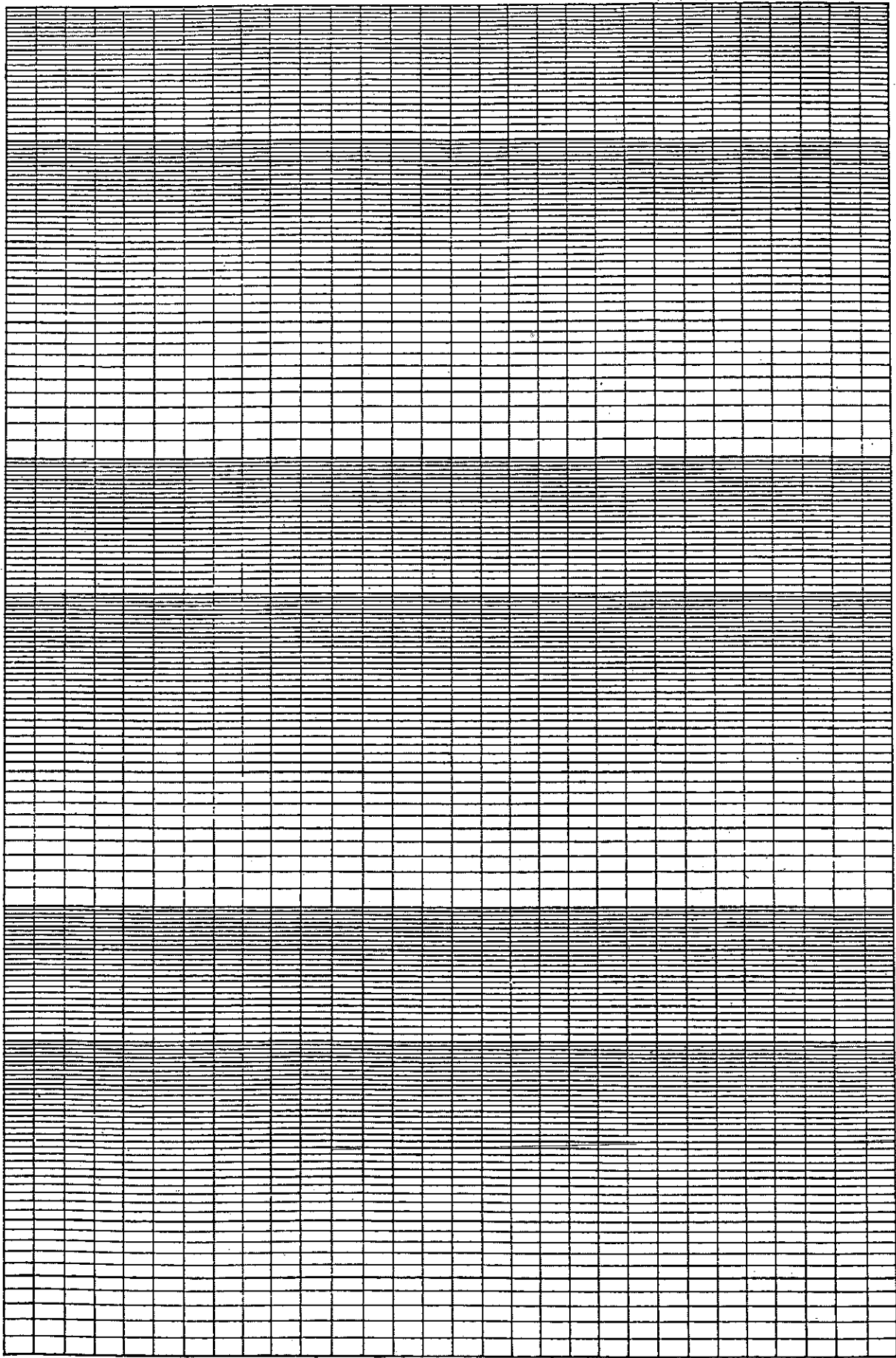


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In[8]:= Export["ElectricField.pdf", %]

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Out[8]= ElectricField.pdf

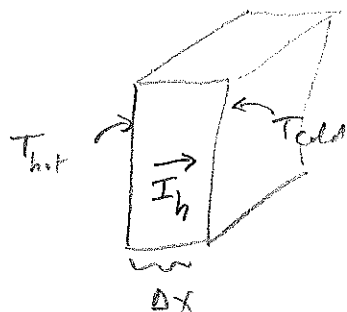


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2. Thermal conduction: heat current (flow of thermal energy) is driven by temperature gradients



Consider an infinitesimal slab whose faces are at different temperatures.
Heat flows across slab.

Define heat current $I_h \equiv \frac{\text{thermal energy flow}}{\text{time}}$

Difference of temperature causes heat flow:

$$I_h = \frac{(\text{temp. difference})(\text{area of slab})}{(\text{thickness})} \quad \left(\begin{array}{l} \text{material} \\ \text{dependent} \\ \text{constant} \end{array} \right)$$

↑
[argument: if double thickness, then ΔT is halved for each half]
[also: insulation!]

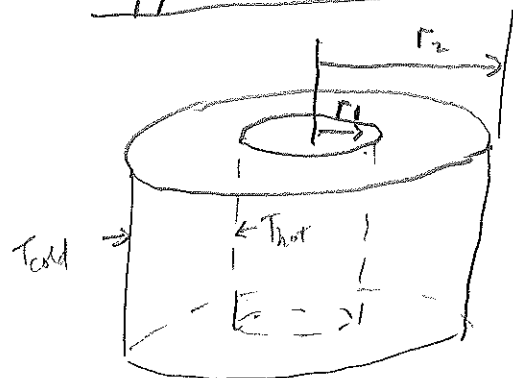
$$I_h = - \frac{\Delta T \cdot A \cdot k}{\Delta x}$$

$$k = \text{thermal conductivity} \left(\frac{\text{J}}{\text{s} \cdot \text{m} \cdot \text{K}} \right) \quad \left\{ \begin{array}{l} \text{silver} \approx 400 \\ \text{glass} \approx 1 \\ \text{air} \approx 0.023 \end{array} \right.$$

Define heat current density $J_h = \frac{I_h}{A}$ [per unit area]

$$J_h = - \frac{k \Delta T}{\Delta x} \xrightarrow{\Delta x \rightarrow 0} -k \frac{dT}{dx} \quad (\text{Fourier's law})$$

Application: insulation around a hot pipe



Assume steady-state (indep. of time)

Energy conservation
 $\Rightarrow$ heat flow in = heat flow out
 (no build up)

Heat flows radially $\Rightarrow$ choose infinitesimal shells as thin cylindrical shells of radius r & thickness Δr .

Area of shell $A = (2\pi r) L$

$$\text{Heat flow through shell } I = -kA \frac{dT}{dr} \\ = -2\pi kL r \frac{dT}{dr}$$

I must indep of t and indep of r , $\therefore$ constant

$$\frac{dT}{dr} = - \frac{I}{2\pi kL r} \quad [\text{separable}]$$

$$T = - \frac{I}{2\pi kL} \log r + C_1$$



Apply boundary conditions:

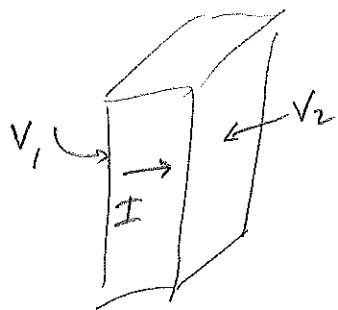
$$T_{\text{hot}} = - \frac{I}{2\pi kL} \log r_1 + C_1$$

$$T_{\text{cold}} = - \frac{I}{2\pi kL} \log r_2 + C_1$$

$$T_{\text{hot}} - T_{\text{cold}} = \frac{I}{2\pi kL} (\log r_2 - \log r_1) = \frac{I}{2\pi kL} \log \left(\frac{r_2}{r_1} \right)$$

$$\text{Heat flow } I = \frac{2\pi kL (T_{\text{hot}} - T_{\text{cold}})}{\log \left(\frac{r_2}{r_1} \right)}$$

3. Electrical conduction: electrical current (flow of charge) driven by voltage gradients



Consider an infinitesimal slab of resistive material whose faces are at different voltages
 $\Rightarrow$ current flows through slab

Define electrical current $I = \frac{\text{charge transfer}}{\text{time}}$

Voltage difference causes current flow

$$I = - \frac{(\Delta V) A \sigma}{\Delta x}$$

$\sigma =$ electrical conductivity

Current density $J = \frac{I}{A}$

x -component $J_x = -\sigma \frac{\partial V}{\partial x} = \sigma E_x$ since $E_x = -\frac{\partial V}{\partial x}$

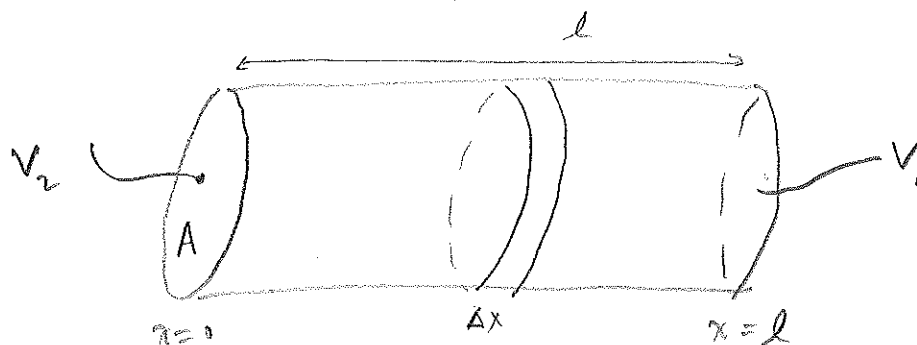
3 dimensions $\vec{J} = \sigma \vec{E}$ (Ohm's law)

$\rho =$ electrical resistivity $\equiv \frac{1}{\sigma}$

$$\vec{E} = \rho \vec{J}$$

$$\begin{cases} \text{Cu: } \rho = 1.7 \times 10^{-8} \Omega \cdot \text{m} \\ \text{Ge } \rho = 0.46 \\ \text{glass } \rho = 10^{10} \end{cases}$$

Application: cylindrical resistor



Assume steady state.

$$I = -\sigma A \frac{dV}{dx}$$

I is independent of x , otherwise charge would build up

$$\frac{dV}{dx} = -\frac{I}{\sigma A}$$

$$V = -\frac{I}{\sigma A} x + C_1$$

Boundary conditions: $V_2 = V(0) = C_1$

$$V_1 = V(l) = -\frac{I}{\sigma A} l + C_1$$

$$V_2 - V_1 = I \left(\frac{l}{\sigma A} \right)$$

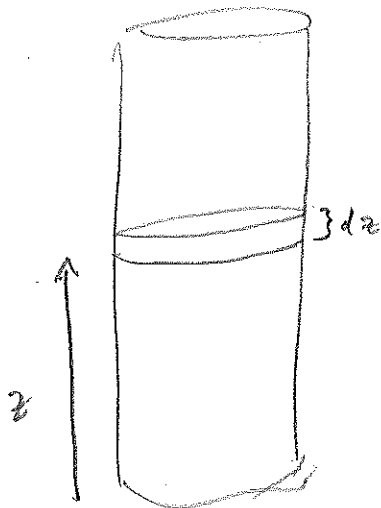
but $\Delta V = IR$

resistance of cylinder is $R = \frac{l}{\sigma A} = \frac{\rho l}{A}$

4. Continuum mechanics

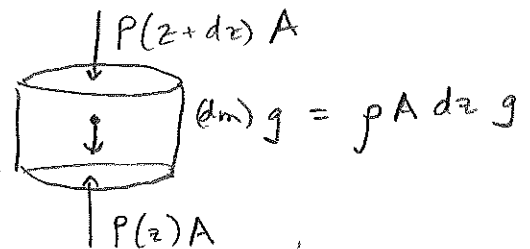
- strategy:
- break system into infinitesimal elements
 - draw free body diagrams for each element
 - apply $\vec{F} = m\vec{a}$ to each element

Static column of fluid:



P = Pressure = force/area, ρ = mass/volume

FBD:



Static $\Rightarrow$ each element in mechanical equilibrium

$$F_{\text{net}}^z = P(z)A - \underbrace{P(z+dz)A}_{P(z) + \frac{dP}{dz} dz + \dots} - \underbrace{(dm)g}_{\rho A dz} = 0$$

$$\Rightarrow \frac{dP}{dz} = -\rho g \quad (\text{hydrostatic eqn.})$$

To proceed, need to know how P and ρ are related (equation of state)

② Assume incompressible fluid.

$\rho = \text{const}$, independent of P and $\therefore z$

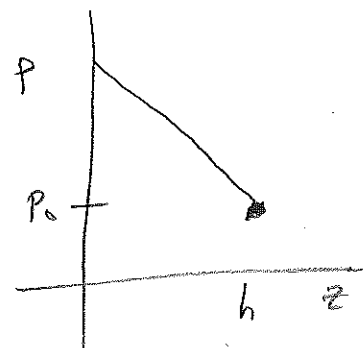
$$\frac{dP}{dz} = -\rho g$$

$$P = -\rho g z + C_1$$

Boundary condition: $P = P_0$ at top of column of fluid $z = h$

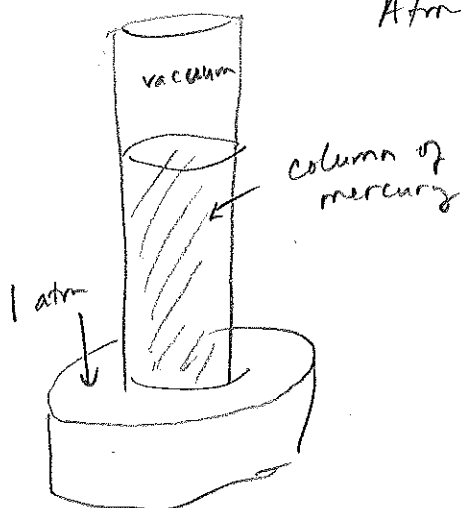
$$P_0 = -\rho g h + C_1$$

$$\Rightarrow P = P_0 + \rho g (h - z)$$



Let's compute atmospheric pressure.

Atmosphere supports 760 mm of mercury



$$\begin{aligned} \underbrace{P(0)}_{1 \text{ atm}} &= \underbrace{P_0}_0 + \underbrace{\rho g h}_{(13.5 \frac{\text{g}}{\text{cm}^3})(9.8 \frac{\text{m}}{\text{s}^2})(0.76 \text{ m})} \\ &= 1.013 \times 10^5 \text{ Pa} \end{aligned}$$

$$(1 \text{ Pascal} = \frac{\text{N}}{\text{m}^2})$$

[pressure increases by 1 atmosphere every 10 m of water]

$$\rho g = (1 \frac{\text{g}}{\text{cm}^3})(9.8 \frac{\text{m}}{\text{s}^2}) = 10 \frac{\text{N}}{\text{m}^3} = \frac{0.1 \text{ atm}}{10 \text{ m}}$$

If fluid is not incompressible, ρ depends on P
(equation of state)

9/10

① ideal gas eqn of state $PV = NRT$

$$\begin{cases} N = \# \text{ moles} \\ R = N_A k_B = 8.31 \frac{\text{J}}{\text{K} \cdot \text{mole}} \end{cases}$$

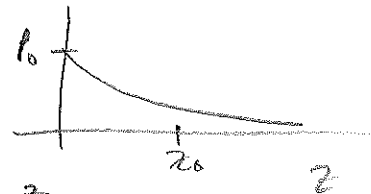
$$\text{mass density } \rho = \frac{mN}{V}, \quad m = \text{mass per mole}$$

$$\Rightarrow P = \rho \frac{RT}{m}$$

$$\therefore \text{hydrostatic eqn becomes } \frac{dP}{dz} = -\left(\frac{mg}{RT}\right) P$$

Assume isothermal atmosphere $\Rightarrow T = \text{const} \Rightarrow 1$ dependent variable

$$P(z) = A e^{-\frac{mgz}{RT}}$$



Apply b.c. $P = P_0$ at $z = 0$

$$P(z) = P_0 e^{-\frac{mgz}{RT}} = P_0 e^{-\frac{z}{z_0}} \quad (\text{Barometric eqn})$$

$$z_0 = \frac{RT}{mg} \approx \frac{(8.31 \frac{\text{J}}{\text{K} \cdot \text{mole}})(300 \text{ K})}{(0.028 \text{ kg/mole})(9.8 \text{ m/s}^2)} \approx 9100 \text{ m} \approx 9 \text{ km}$$

$$N_2: 14 \times 2$$

fall off slower for lighter species (H, He)