

Nuclear models

~~EN-6~~
(10) NM-1
(nuclear models)

[observe that more abundant nuclei generally have smaller mass excess $\rightarrow$ greater binding energy $\Rightarrow$ more stable]

neutral atom $E_{\text{rest}} = Z m_H c^2 + N m_n c^2 + \underbrace{\left(\begin{smallmatrix} \text{internal} \\ \text{kinetic} \\ \text{energy} \end{smallmatrix} \right)}_{>0} + \underbrace{\left(\begin{smallmatrix} \text{attractive} \\ \text{potential} \\ \text{energy} \end{smallmatrix} \right)}_{<0}$

total must be negative, else nucleus would be unstable to decay into p & n

$$= -B$$

binding energy $B = Z m_H c^2 + N m_n c^2 - E_{\text{rest}}$

$$= Z (938.8 \text{ MeV}) + N (939.6 \text{ MeV}) - A (931.5 \text{ MeV}) - \Delta$$

$$= Z (7.29 \text{ MeV}) + N (8.07 \text{ MeV}) - \Delta$$

$$= A (7.68 \text{ MeV}) + (A - 2Z) (0.39 \text{ MeV}) - \Delta$$

$$Z = \frac{A}{2} - \frac{1}{2}(A - 2Z)$$

$$N = \frac{A}{2} + \frac{1}{2}(A - 2Z)$$

If $Z = N = \frac{1}{2}A$ and $\Delta \approx 0$, then

$$B \approx (7.7 \text{ MeV}) A$$

i.e. binding energy proportional to A

$$\frac{B}{A} \sim 8 \text{ MeV} \quad [\text{look at curve} \Rightarrow]$$

[Actually $\frac{B}{A}$ higher for middle nuclei, less at ends $\Rightarrow$ Explain!]
"Curve of binding energy"

[use table of Δ to compute & plot $\frac{B}{A}$ for most abundant nuclei w/ $A < 32$]

$$\frac{E_{\text{rest}}}{Z m_H c^2 + N m_n c^2} = 1 - \frac{B}{Z m_H c^2 + N m_n c^2} \approx 1 - \frac{B}{A (940 \text{ MeV})} \approx 1 - 0.01$$

Nuclei are $\sim 1\%$ less massive than sum of constituents

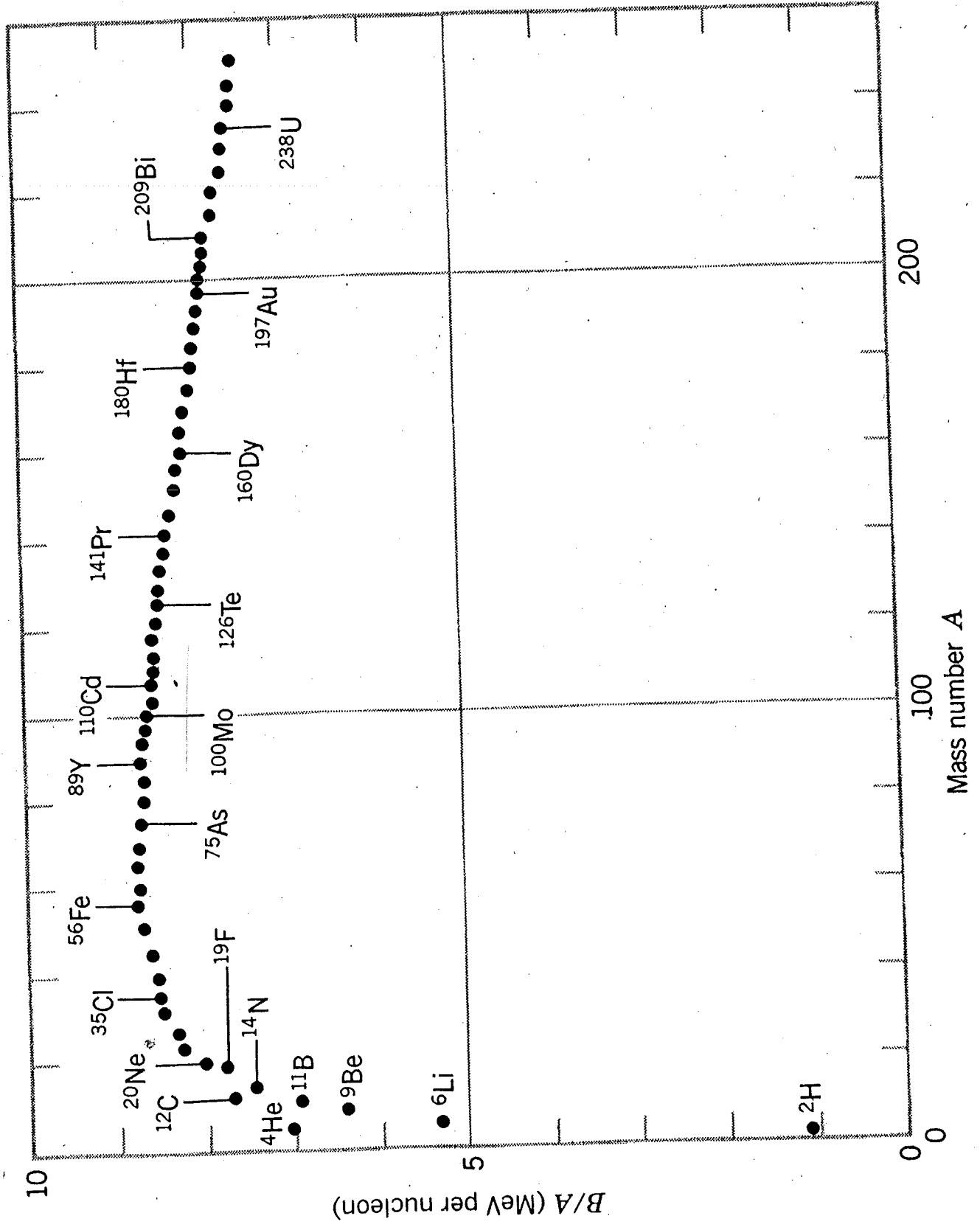


Figure 3.16 The binding energy per nucleon.

B/A for most abundant isotopes

⊙ = even-even nuclides

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$\frac{B}{A}$

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MILLIMETER

9

8

7

6

5

4

3

2

1

5

10

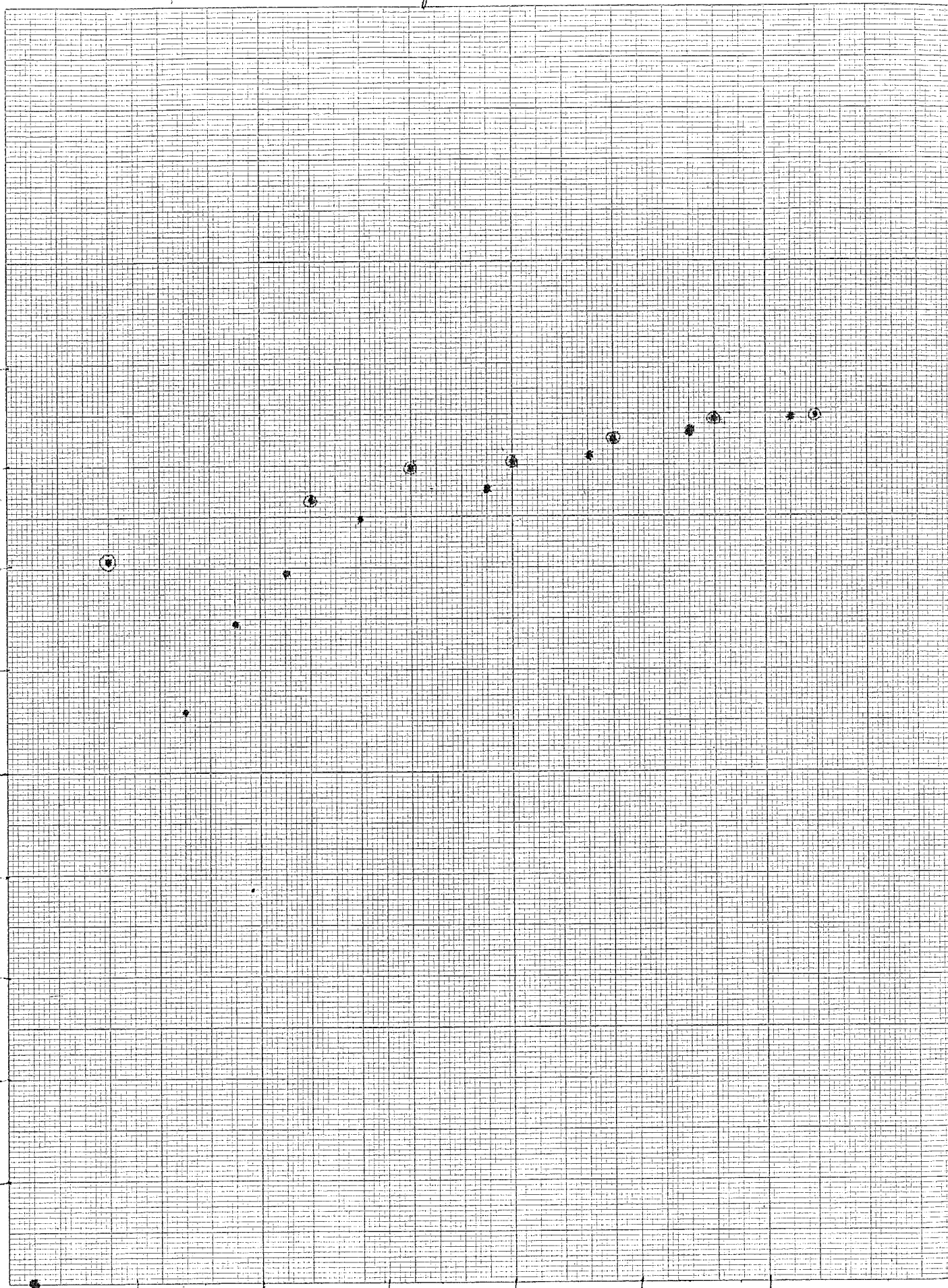
15

20

25

30

A



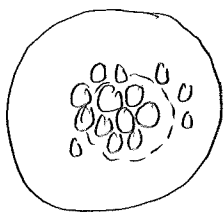
[Try to understand binding energy]

[Force primarily responsible for existence of nucleus is

strong nuclear force

• acts between pairs of nucleons (pp, pn, nn)

• short range, range $\ll$ size of nucleus



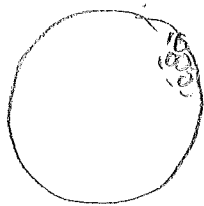
Suppose each nucleon acts on r others for some r

total # of pair ("bonds") is $\frac{1}{2}rA$

Potential energy = - (# pairs) (energy of "bond")

$$\Rightarrow B \approx c_1 A \quad (\text{volume term})$$

Then $\frac{B}{A} \approx c_1$ explains why binding energy per nucleon is $\approx$ const
(due to short range nature of force, else $B \sim A^2$)



Nucleus on surface "lose" fewer "bonds"

"missing bonds" $\propto$ # of nucleons on surface $\sim 4\pi R^2 \sim A^{2/3}$

$$B \approx c_1 A - c_2 A^{2/3} \quad (\text{surface term})$$

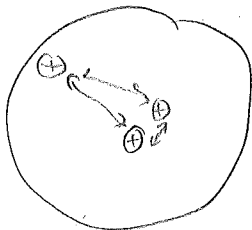
Then $\frac{B}{A} \approx c_1 - \frac{c_2}{A^{1/3}}$ explains why curve dips down for small A

Surface term causes nucleus to minimize surface area $\Rightarrow$ sphere

Liquid drop model of nucleus

Electromagnetic force

- acts between pairs of protons
- long range : affects all protons in nucleus



- repulsive, ∇U , reduces binding energy

Treat protons as uniformly distributed charge $Q = Ze$

Potential energy of a sphere of charge $U = \frac{3}{5} \frac{KQ^2}{R} = \underbrace{\left(\frac{3Ke^2}{5r_0} \right)}_{c_3} \frac{Z^2}{A^{1/3}}$

$$B \approx c_1 A - c_2 A^{2/3} - c_3 \frac{Z^2}{A^{1/3}}$$

$$\frac{B}{A} \approx c_1 - \frac{c_2}{A^{1/3}} - c_3 \frac{Z^2}{A^{4/3}}$$

explains why $\frac{B}{A}$ decreases for large Z, A + why nuclei eventually become unstable

[electromagnetic force eventually wins because it acts on all pairs of protons whereas strong only acts on neighbouring nucleons]

$$\text{Estimate } c_3 = \frac{3}{5} \frac{Ke^2}{r_0} = \frac{3}{5} \frac{(8.99 \times 10^9 \text{ J}\cdot\text{m}/\text{C}^2) (1.6 \times 10^{-19} \text{ C})^2}{(1.2 \times 10^{-15} \text{ m})} =$$

$$= 1.15 \times 10^{-13} \text{ J}$$

$$= 0.72 \text{ MeV}$$

[eg Krane, p. 67]

$$\left[\frac{3}{5} \frac{Ke^2}{\hbar c} \frac{\hbar c}{r_0} = \frac{3}{5} \alpha \frac{197 \text{ MeV}\cdot\text{fm}}{1.2 \text{ fm}} \right]$$

EN 9

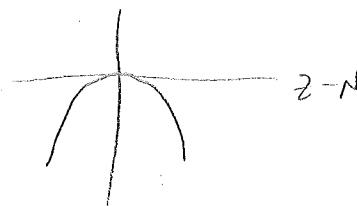
NMY

Symmetry term (origin quantum mechanical)

acts to keep # protons + neutrons approx equal

$$\sim -\frac{(Z-N)^2}{A}$$

$$\sim -\frac{(A-2Z)^2}{A}$$



Semi-empirical binding energy formula (Bethe-Weizsäcker) (fits data well for $A \geq 15$)

$$B = \underbrace{c_1 A}_{\text{strong}} - \underbrace{c_2 A^{2/3}}_{\text{electromagnetic}} - \underbrace{c_3 \frac{Z^2}{A^{1/3}}}_{\text{quantum effects}} - \underbrace{c_4 \frac{(A-2Z)^2}{A}}_{\text{quantum effects}} + \delta_{\text{pair}}$$

[Krauss]

$$c_1 = 15.5 \text{ MeV}$$

$$c_2 = 16.8 \text{ MeV}$$

$$c_3 = 0.72 \text{ MeV}$$

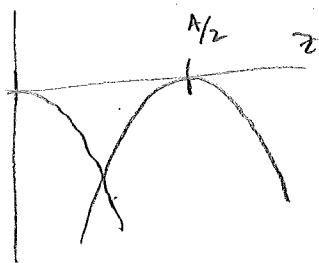
$$c_4 = 23 \text{ MeV}$$

$$c_5 = 34 \text{ MeV}$$

$$\delta_{\text{pair}} = \begin{cases} c_5 A^{-3/4} & Z, N \text{ even} \\ 0 & Z, N \text{ odd} \\ -c_5 A^{-3/4} & Z, N \text{ odd} \end{cases}$$

3rd + 4th terms determine the line of stability $\Rightarrow$

$$B = c_1 A - c_2 A^{2/3} - c_3 \frac{Z^2}{A^{1/3}} - c_4 \frac{(A-2Z)^2}{A} + \delta_{pair}$$



favors $Z \ll N$

more effective at large A



heavy nuclei have fewer protons than neutrons

favors $Z = \frac{A}{2}$

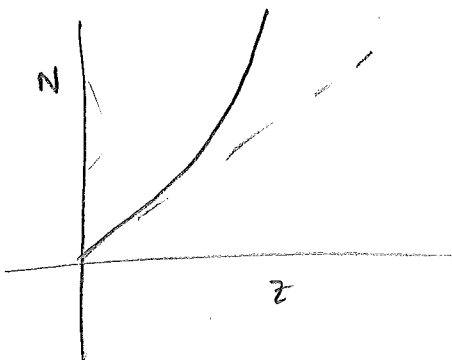
$\approx Z = N$

more effective at small A



light nuclei have $Z \approx N$

${}^4\text{He}, {}^{12}\text{C}, {}^{16}\text{O}$

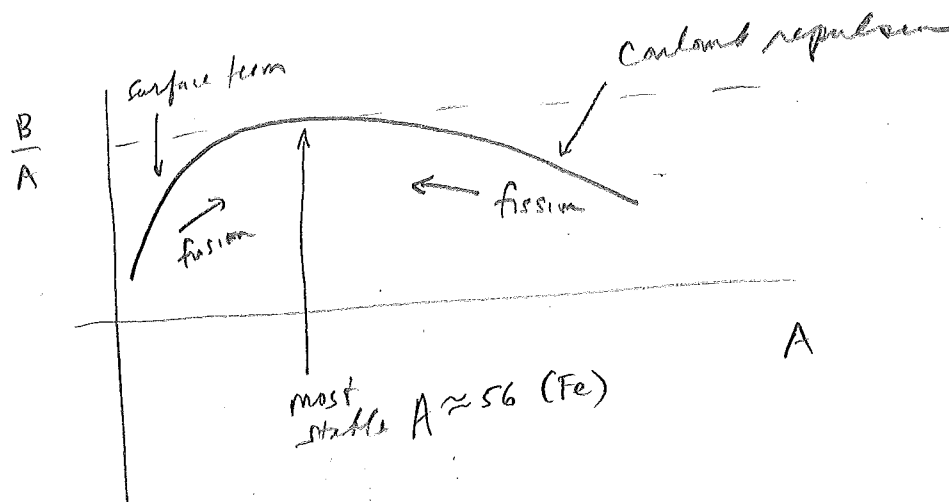


Binding energy per nucleon

$$\frac{B}{A} = c_1 - c_2 A^{-1/3} - c_3 \frac{Z^2}{A^{1/3}} - c_4 \frac{(A-2Z)^2}{A^2} + \frac{\delta_{pair}}{A}$$

assume $Z \approx N \approx \frac{A}{2}$

$$\frac{B}{A} = c_1 - \frac{c_2}{A^{1/3}} - c_3 A^{2/3}$$



$$X \rightarrow Y + Z \quad \text{iff} \quad m_X > m_Y + m_Z$$

A nucleus is stable only if its mass is less than the sum of masses of nuclei into which it could decay

[First 2 terms of semi-empirical explained by liquid drop.

3rd term $\rightarrow$ Coulomb repulsion.

4th term $\rightarrow$ symmetry term } OM effects.

5th term $\rightarrow$ pairing term

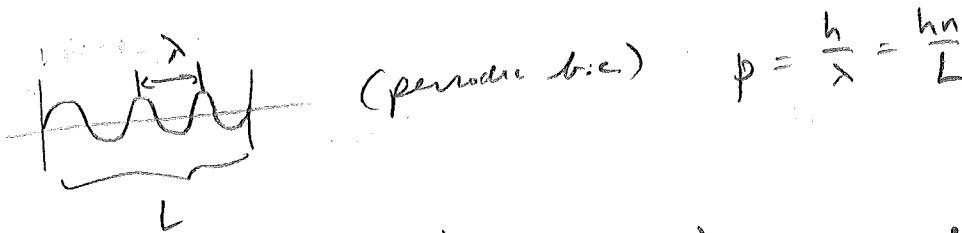
How understand symmetry term?]

Fermi gas model of nucleus

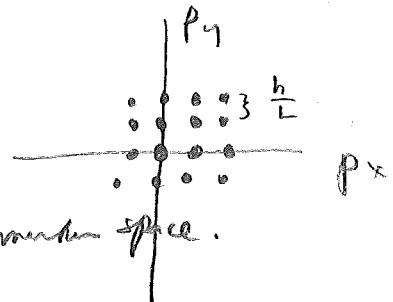
treat protons + neutrons as non-interacting fermions confined to V .

For simplicity, cube of side $L = V^{1/3}$

p+n have de Broglie wavelength $\lambda = \frac{h}{p}$ where $\lambda = \frac{L}{n}$, $n \in \mathbb{Z}$



3 dimensions, $\vec{p} = \frac{h}{L}(n_x, n_y, n_z)$



Each state takes up volume $(\frac{h}{L})^3$ in momentum space.

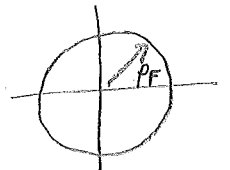
Identical fermions occupy different states, filling out a sphere of radius p_F

Total # of $\frac{1}{2}$ (spin- $\frac{1}{2}$) fermions $N = 2 \left(\frac{L}{h}\right)^3 \int dp_x dp_y dp_z$

$$= 2 \frac{V}{h^3} \int_0^{p_F} 4\pi p^2 dp = \frac{8\pi V p_F^3}{3h^3} = \frac{V p_F^3}{3\pi^2 h^3}$$

$$\Rightarrow p_F = \left(\frac{3\pi^2 N h^3}{V} \right)^{1/3} = \left(\frac{9\pi}{4} \frac{N}{A} \right)^{1/3} \frac{h}{m_0}$$

Recall $V = \left(\frac{4}{3}\pi r_0^3\right) A$



Is the Fermi gas relativistic or nonrelativistic?

$$E = \sqrt{(c p_F)^2 + (m c^2)^2}$$

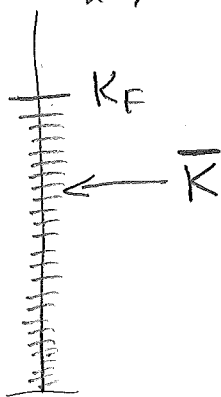
$$\begin{array}{ll} c p_F \ll m c^2 & \text{nonrel} \\ c p_F \gg m c^2 & \text{ultra rel} \end{array}$$

Estimate $c p_F \sim \frac{\hbar c}{r_0} = \frac{(197 \text{ MeV fm})}{(1.2 \text{ fm})} \sim 170 \text{ MeV} \ll 940 \text{ MeV}$

Treat it as nonrel (all momenta are $\leq p_F$)

$$K = \frac{p^2}{2m} \Rightarrow K_F = \frac{p_F^2}{2m} = \frac{\hbar^2}{2m r_0^2} \left(\frac{q_0 N}{4A} \right)^{2/3}$$

Show that the average kinetic energy of a nonrelativistic fermi gas is $\bar{K} = \frac{3}{5} K_F$ [HW]



Therefore total kinetic energy of N identical (spin- $\frac{1}{2}$) fermions "

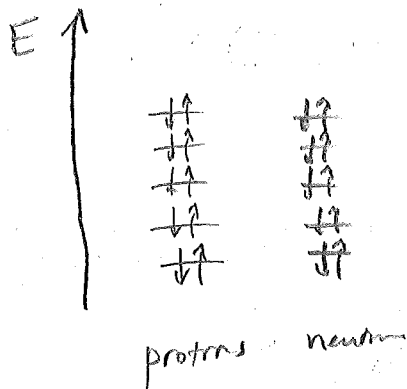
$$K_{\text{tot}} = \left(\frac{3}{5} K_F \right) N = \frac{3 \hbar^2}{10 m r_0^2} \left(\frac{q_0}{4} \right)^{2/3} \left(\frac{N}{A} \right)^{5/3} \cdot A$$

$$\frac{\hbar^2}{m r_0^2} = \frac{(\hbar c)^2}{(m c^2) r_0^2} = \frac{(197 \text{ MeV fm})^2}{(940 \text{ MeV})(1.2 \text{ fm})^2} \approx 29 \text{ MeV}$$

Include protons and neutrons

$$K_{\text{tot}} = \frac{3}{10} \left(\frac{q_0}{4} \right)^{2/3} \left[\left(\frac{Z}{A} \right)^{5/3} + \left(\frac{A-Z}{A} \right)^{5/3} \right] (29 \text{ MeV}) \cdot A$$

First, assume $Z = N = \frac{1}{2} A$

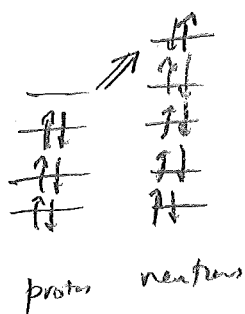


$$K_{\text{tot}} = \underbrace{\frac{3}{10} \left(\frac{90}{4} \right)^{2/3} \left[\left(\frac{1}{2} \right)^{5/3} + \left(\frac{1}{2} \right)^{5/3} \right]}_{0.7} (29 \text{ MeV}) A$$

$$= (20 \text{ MeV}) A$$

This contributes to the volume term
 implies that the potential energy per nucleon
 is $\sim -35 \text{ MeV}$

If $Z \neq N$, kinetic energy increases

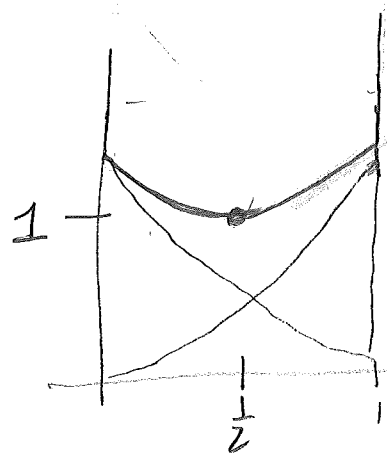


Estimate how much:

Let $x = \frac{Z}{A}$ then

$$K = (20 \text{ MeV}) A \cdot \frac{f(x)}{f(\frac{1}{2})}$$

where $f(x) = x^{5/3} + (1-x)^{5/3}$



$$\frac{f(x)}{f(\frac{1}{2})} = 1 + \# \left(x - \frac{1}{2} \right)^2 + \dots$$

↑
[Calculate this & use it to estimate c_y in semi-empirical, [given: $\# = \frac{20}{9}$]]

[Hint: it doesn't agree too well but is right order of magnitude]

[$c_y = 11 \text{ MeV}$ whereas actual $c_y = 29 \text{ MeV}$]

Kinetic energy of nucleus

$$\vec{k} = \frac{2\pi}{L} \vec{n}$$

~~4\pi n^3_{max}~~

$$\frac{4\pi}{3} n^3_{max} = \frac{N}{2}$$

$$E_F = \frac{\hbar^2 k^2}{2m} = \frac{\hbar^2}{2m} \left(\frac{2\pi}{L} \right)^2 \left(\frac{3N}{8\pi} \right)^{2/3} = \frac{\hbar^2}{2m} \left(\frac{3\pi^2 N}{V} \right)^{2/3}$$

$$U = \frac{3}{5} N E_F$$

$$r_0 = 1.2 \text{ fm (Semi)} \\ = 1.5 \text{ fm (Fermi)}$$

For a nucleus: $V = A V_0$, $V_0 = \frac{4}{3} \pi r_0^3$

$$E_F = \frac{\hbar^2 c^2}{2m c^2 r_0^2} \left(\frac{\pi}{4} \right)^{2/3} \left(\frac{N}{A} \right)^{2/3}$$

$$\begin{matrix} \nearrow \\ C_0 \end{matrix} \quad \begin{matrix} 53 \text{ MeV (Semi)} \\ 34 \text{ MeV (Fermi)} \end{matrix}$$

$$\rightarrow N_p = N_n = \frac{A}{2} \Rightarrow E_F \approx 21 \text{ MeV (F)}$$

$$\left\{ \begin{array}{l} \text{if only neutrons} \\ E_F = \begin{cases} 53 \\ 34 \end{cases} \end{array} \right.$$

$$U_{nuc} = \frac{3}{5} N_p C_0 \left(\frac{N_p}{A} \right)^{2/3} + \frac{3}{5} N_n C_0 \left(\frac{N_n}{A} \right)^{2/3}$$

$$= \frac{3}{5} \frac{C_0}{A^{2/3}} \left[Z^{5/3} + (A-Z)^{5/3} \right] = f(Z, A)$$

$$p = Z - \frac{A}{2}$$

$$f(Z, A) = \left(\frac{A}{2} + p \right)^{5/3} + \left(\frac{A}{2} - p \right)^{5/3}$$

$$= \frac{A^{5/3}}{2^{3/2}} \left[1 + \frac{20}{9} \frac{p^2}{A^2} \right] = (.63) A^{5/3} + (.35) A^{-1/3} (A-2Z)^2$$

$$U_{Nuc} = (.38) C_0 A + (.21) C_0 \frac{(A-2Z)^2}{A}$$

$\begin{matrix} 20 \text{ MeV (Semi)} \\ 13 \text{ MeV (Fermi)} \end{matrix} \quad \begin{matrix} 11 \text{ MeV (Semi)} \\ 7 \text{ MeV (Fermi)} \end{matrix}$

If all neutrons, $(A=N_n, Z=0)$

$$U_{Nuc} = \frac{3}{5} C_0 A = \begin{cases} 32 \text{ MeV} \\ 20 \text{ MeV} \end{cases} A$$

$$\begin{matrix} 17 \text{ MeV} \\ 7 \text{ MeV} \\ 10 \text{ MeV} \end{matrix}$$