

(2019)

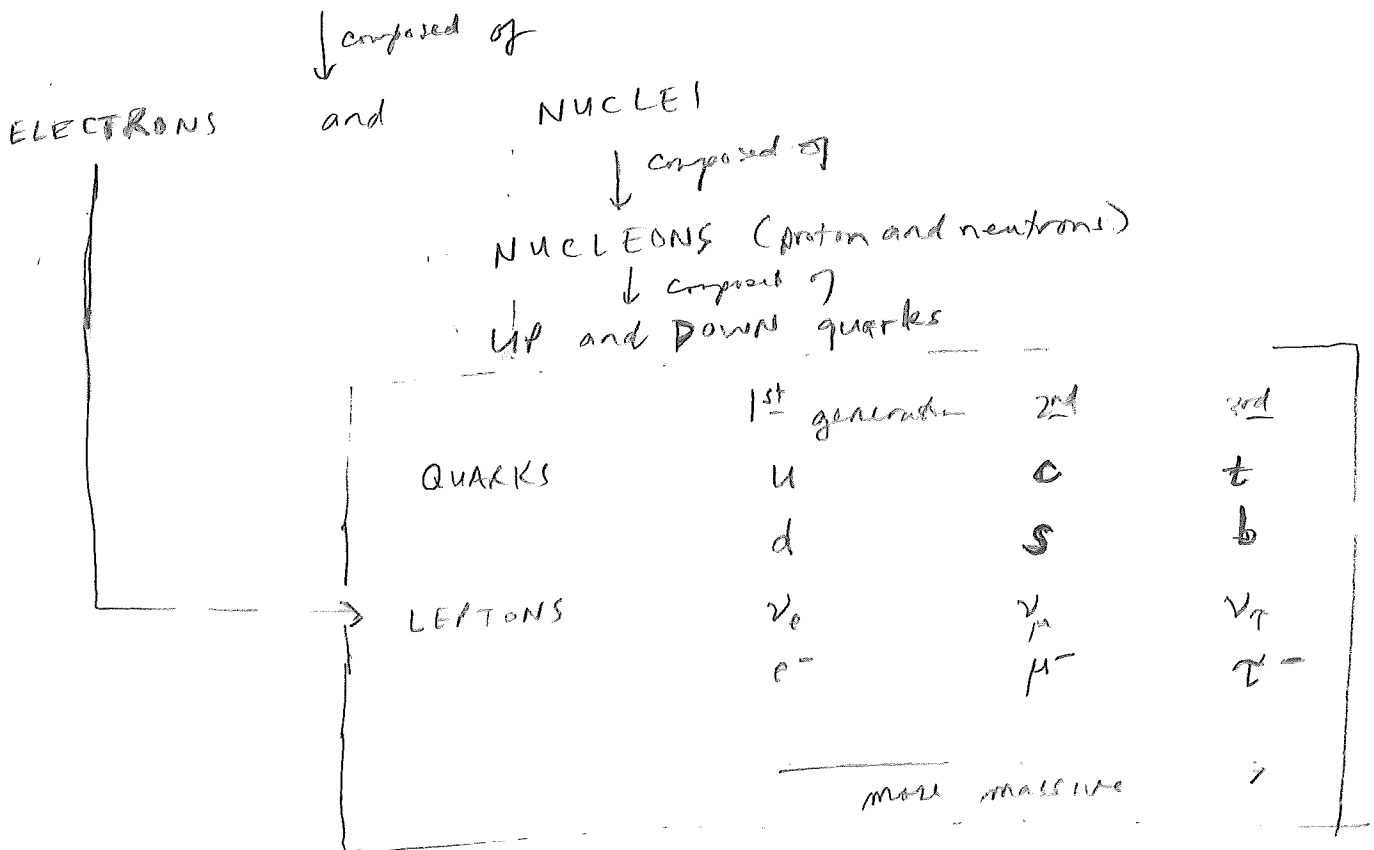
PHYS 2260 Nuclear & Particle Physics

pi-1
[particles & interactions]

WHAT ARE THE FUNDAMENTAL CONSTITUENTS OF NATURE?

[What is everything made out of?
From ancient times, people have asked this question.
By the end of 19th century, most chemists & physicists concluded

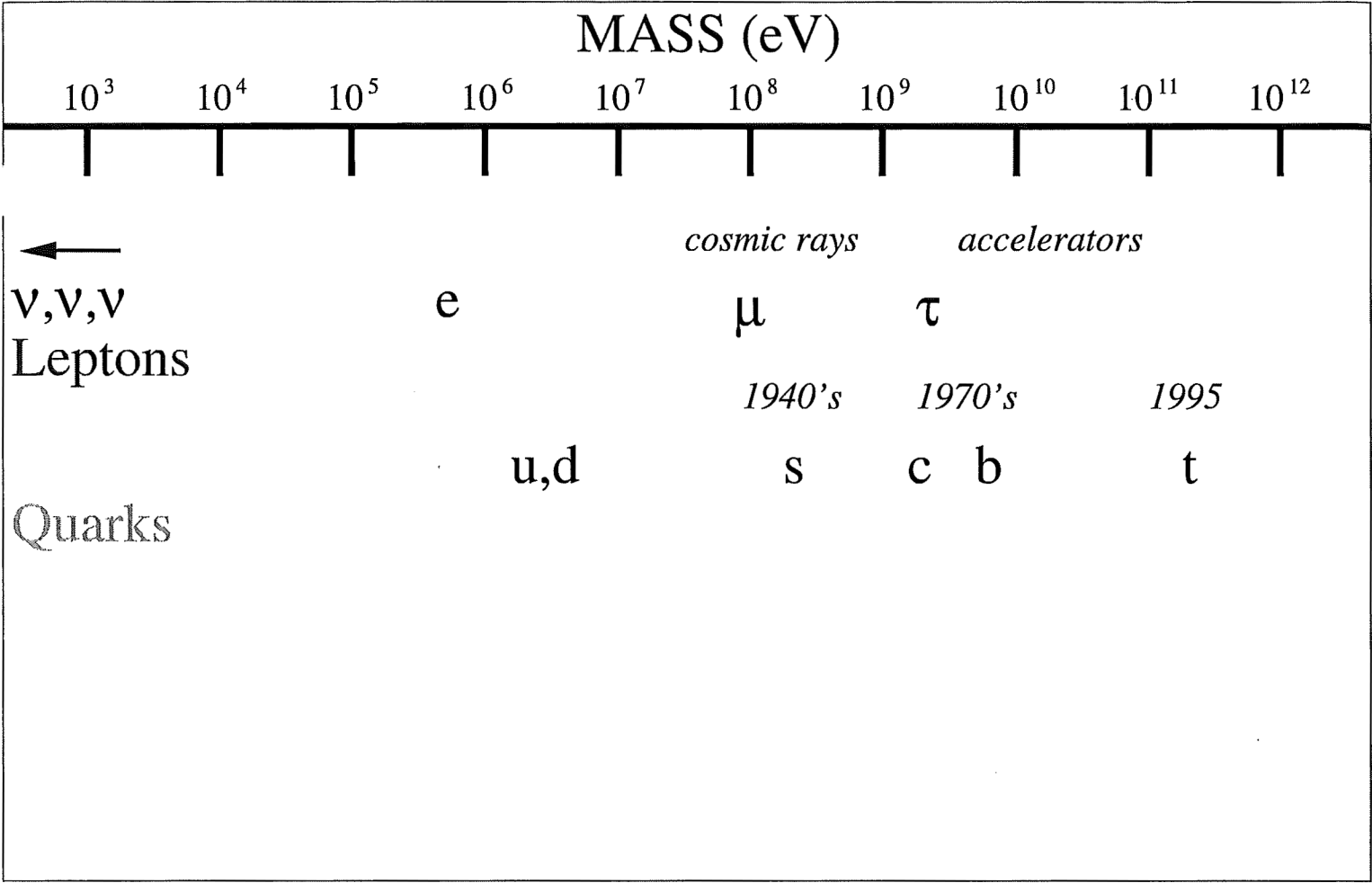
1. Atoms (≈ 100 different types)



Neutrinos initially believed to be massless.
They actually have small (but yet unknown) masses.

There 6 Quarks & 6 leptons are believed to be the most fundamental constituents
but...

more generations? $\rightarrow$ not unless $m_\nu > 50$ times mass of proton
(only top quark is that heavy)
Subquarks & subleptons? $\rightarrow$ no evidence
Strings?



This is only half the story

pi-2

[The theoretical framework which describes these particles is]

quantum field theory (QFT), merge of QM and special relativity

QFT $\Rightarrow$ for each type of particle, there is a corresponding type of antiparticle, w/ same mass + opposite charge

<u>antiquarks</u>	$\bar{u}$	$\bar{c}$	$\bar{t}$
	$\bar{d}$	$\bar{s}$	$\bar{b}$
anti-lep/ms	$\bar{\nu}_e$	$\bar{\nu}_\mu$	$\bar{\nu}_\tau$
	e^+	μ^+	τ^+
	$\uparrow$		
	positron		

[use bars on all
xc charged leptons]

Antiparticles are generally much less abundant
(despite the theoretical symmetry between pds + antipds)

[Why?]

Interactions ("forces")

- 1) gravity [oldest known, + weakest]
- 2) electromagnetism
- 3) weak nuclear force
- 4) strong nuclear force

[other fundamental interactions? "fifth force expts"
none known so far, but likely in GUTs; much weaker]

Interactions cause:

- 1) bound states
- 2) elastic scattering [particle identities unchanged] Rutherford
- 3) "rearrangements" [chemical reactions]

But also

4) decays: $n \rightarrow p e^- \bar{\nu}_e$

- 5) inelastic scattering: change of particle identity
or production of new particles
by cosmic rays or
in particle accelerators

Interactions are complicated, but
it is observed that some quantities are conserved
(some before & after)

total mass is not conserved

Conserved quantities

- energy
- momentum
- angular momentum
- electric charge

believed to be absolute, conserved by
all known & unknown interactions

- baryon # (quarks/antiquarks have baryon # $\frac{1}{3}/-\frac{1}{3}$; could... $\therefore$ q's #)
- lepton # (= #leptons)

conserved by all known forces (standard model)
but perhaps not by GUTs.
conservation of q's, cons. of leptons

Partially conserved

... conserved by strong nuclear force & EM
but not by weak nuclear force

- strangeness
- isospin
- C = charge conjugation
- P = parity (space reversal)
- T = time reversed symmetry

Most elementary particles are unstable.

more massive particle inevitably decay into
less massive particles unless
forbidden by a conservation law

All charged 2nd & 3rd gen. pds decay to 1st gen

u, d, ν_e, e^- ; ν_μ, ν_τ

More massive particles can be created in collisions
(eg particle accelerators)

Skip this

π^-

[We characterize different types of particles (both fundamental + composite) by their properties]

PROPERTIES OF PARTICLES

- MASS
- CHARGE, ELECTRIC + OTHERWISE (eg. quarks also have color charge)
- OTHER CONSERVED (OR ALMOST CONSERVED) ADDITIVE QUANTUM NUMBERS (eg. baryon no., strangeness)
- MULTIPLICATIVE QUANTUM NUMBERS
 - parity P
 - charge conjugation symmetry C
 - time reversal invariance T
- SPIN (ie intrinsic angular momentum)
- MAGNETIC DIPOLE MOMENT
- ISOSPIN, and OTHER FLAVOR QUANTUM NUMBERS
- MEAN LIFE and DECAY PRODUCTS, if unstable

These properties are measured experimentally and
compiled in the Annual Particle Physics Booklet
[hand out; put up on screen]

$\rightarrow \mu^-$

[We'll try to understand and hopefully calculate these properties from our theoretical models]

skip this

pi-5

How do we study elementary particles + learn about their properties?

By throwing things at them + seeing what happens

- sometimes gets deflected or bounces back
- sometimes gets absorbed
- if thrown hard enough, produces new particles

Scattering experiments (target and probe)

Probes

• visible light

[called "looking at something"]

gases scatter $\sim \frac{1}{\lambda^4} \Rightarrow$ Rayleigh scattering

• X-rays

[study crystal structure, DNA]

when scattered from free electrons, $\Rightarrow$ Compton scattering
scattered X-rays have slightly longer λ

• α -particles

scatters off nuclei $\rightarrow$ both positively charged $\Rightarrow$ Rutherford scattering

induces nuclear reactions

• neutrons

- no Coulomb barrier
- absorbed to create

new isotopes or induce fission

• cosmic rays (mostly protons)

create new elementary particles

• accelerators (charged particles: p or e^-)

van de Graaff

cyclotrons

modern particle accelerators (LHC)

① GRAVITY

• long range (inverse square) $F = \frac{Gm_1m_2}{r^2}$

[acts on solar systems, galaxies]

• acts on anything w/ energy (not just mass)
[e.g. bending of light]

• always attractive

[binds hydrogen into stars, stars into galaxies]

• very weak on microscopic scales

[but cumulative so dominant on macroscopic scale]

• ignorable for elementary particles

② ELECTRO. MAGNETISM

• long range (inverse square) $F = \frac{k q_1 q_2}{r^2} = \frac{1}{4\pi\epsilon_0} \frac{q_1 q_2}{r^2}$

[EM waves for distant galaxies]

• acts on all electrically charged particles
(also on neutral particles w/ magnetic moments)

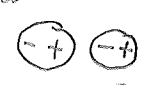
• attraction between opposite charges

[binds negatively charged electrons and positively charged nuclei into (neutral) atoms]

• balance of charges ^{largely} neutralizes Coulomb force on macro scale

• residual forces on neutral objects

binds atoms into molecules (DM plays a role)



Compare electrical + gravitational force on 2 protons

$$K_e^2 = (8.99 \times 10^9 \frac{N \cdot m^2}{C^2}) (1.6 \times 10^{-19} C)^2 = 2.3 \times 10^{-28} N \cdot m^2$$

$$G_m^2 = (6.7 \times 10^{-11} \frac{N \cdot m^2}{kg^2}) (1.7 \times 10^{-27} kg)^2 = 1.9 \times 10^{-64} N \cdot m^2$$

factor of 10^{36}

[force between protons is repulsive EM > gravit,!]
 so need $\downarrow$

p. 8

③ strong nuclear force

• short range, of order $\underline{1 \text{ fm}} \sim 10^{-15} \text{ m}$
 femtometer, or fermi

• acts on particles that have "color charge" (R, G, B)
 e.g. quark or antiquark, ($\bar{R}, \bar{G}, \bar{B}$)
 not leptons

• binds quarks & antiquarks ^(of opposite colors) into mesons ($q\bar{q}$) e.g. $R\bar{R}$
 $\uparrow$
 baryon # = 0

e.g. $\pi^+ = u\bar{d}$
 [many others, all unstable]

• binds triplets of quarks ^(one of each color) into baryons (qqq) ^{RGB} baryon # = 1
 or antiquarks into antibaryons ($\bar{q}\bar{q}\bar{q}$)

e.g. proton $p = uud$
 neutron $n = udd$

[many others, all unstable]

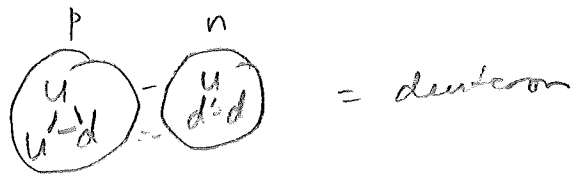
• mesons & baryons are "color neutral" or "color singlets"
 hadrons $\left[\begin{array}{l} \text{technically } \frac{1}{3}(RR\bar{R} + G\bar{G} + B\bar{B}) \\ \frac{1}{6}(RGB - GRB + GBR - BGR + BRG - RBG) \end{array} \right]$
 • other color neutral states may exist, e.g. pentaquarks: $qqqq\bar{q}$

(not qq , or $q\bar{q}q$)

Confinement postulate: color force is so strong that it is believed that only color neutral particles can exist in isolation [no free quarks detected]

→ only leptons, mesons, baryons are observable
allied hadrons

Nucleons (p, n) are color neutral
but the (much weaker) residual color force
can bind them into nuclei



but short range of color force limits the size of nuclei

④ weak nuclear force

- short range
- acts on quarks & leptons
- does not bind particles into composite objects
but is responsible for β -decay of nuclei.

• intergenerational decay ($2 \rightarrow 1$, $3 \rightarrow 2$, $3 \rightarrow 1$)
always occurs as a result of weak interaction

Strength of an interaction determine

- ① binding energy of composite
- ② rate of decay, i.e. mean life τ
- ③ scattering rate, i.e. cross section σ

[interactions $\Rightarrow$ 7 more particles]

EM force mediated by fields $\vec{E}, \vec{B}$

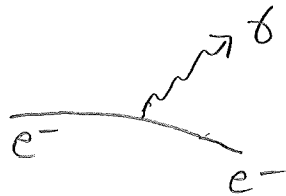
which obey Maxwell eqns

$\Rightarrow$ wave-like solutions that travel at c .

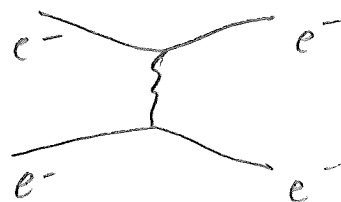
Quantize the EM fields (QED = quantum electrodynamics)

$\Rightarrow$ photons: massless particles $\propto$
($E = cp, v = c$)

Accelerating charges emit photons (bremsstrahlung)



Force between charges caused by
exchange of virtual photons



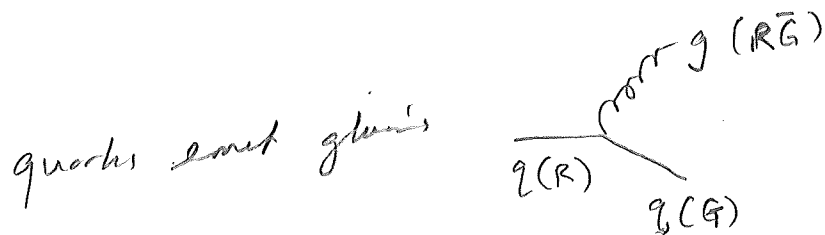
[Feynman
diagram]

strong nuclear force mediated by $SU(3)$ Yang-Mills field
 which obeys Yang-Mills eqs (nonlinear generalization
 of Maxwell)

Quantize Ym fields (QCD - quantum chromodynamics)

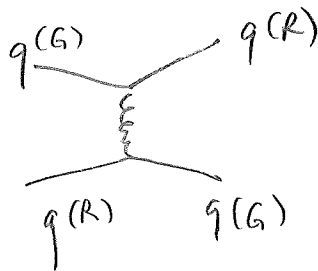
$\Rightarrow$ gluons: massless particles g

8 types:



$$8 = 3 \times 3 - 1$$

exchange of virtual gluons causes the color force



Technically there are 9

$$\begin{pmatrix} R \\ G \\ B \end{pmatrix} \begin{pmatrix} \bar{R} & \bar{G} & \bar{B} \end{pmatrix}$$

$$\begin{pmatrix} R\bar{R} & R\bar{G} & R\bar{B} \\ G\bar{R} & G\bar{G} & G\bar{B} \\ B\bar{R} & B\bar{G} & B\bar{B} \end{pmatrix}$$

but remove color
 singlets

$$\frac{1}{3}(R\bar{R} + G\bar{G} + B\bar{B})$$

weak nuclear force mediated by $SU(2)$ Yang-Mills field

Quantize fields (Glashow Weinberg Salam theory)

$$\Rightarrow W^+, W^-, Z^0 \quad 3 = 2 \times 2 - 1$$

Unlike γ, g , these particles are very massive ($\sim 100 m_p$)

Mass results from SSB (spontaneous symmetry breaking)
caused by Higgs field

Quantize $\Rightarrow$ Higgs boson (also massive, $\sim 130 m_p$)

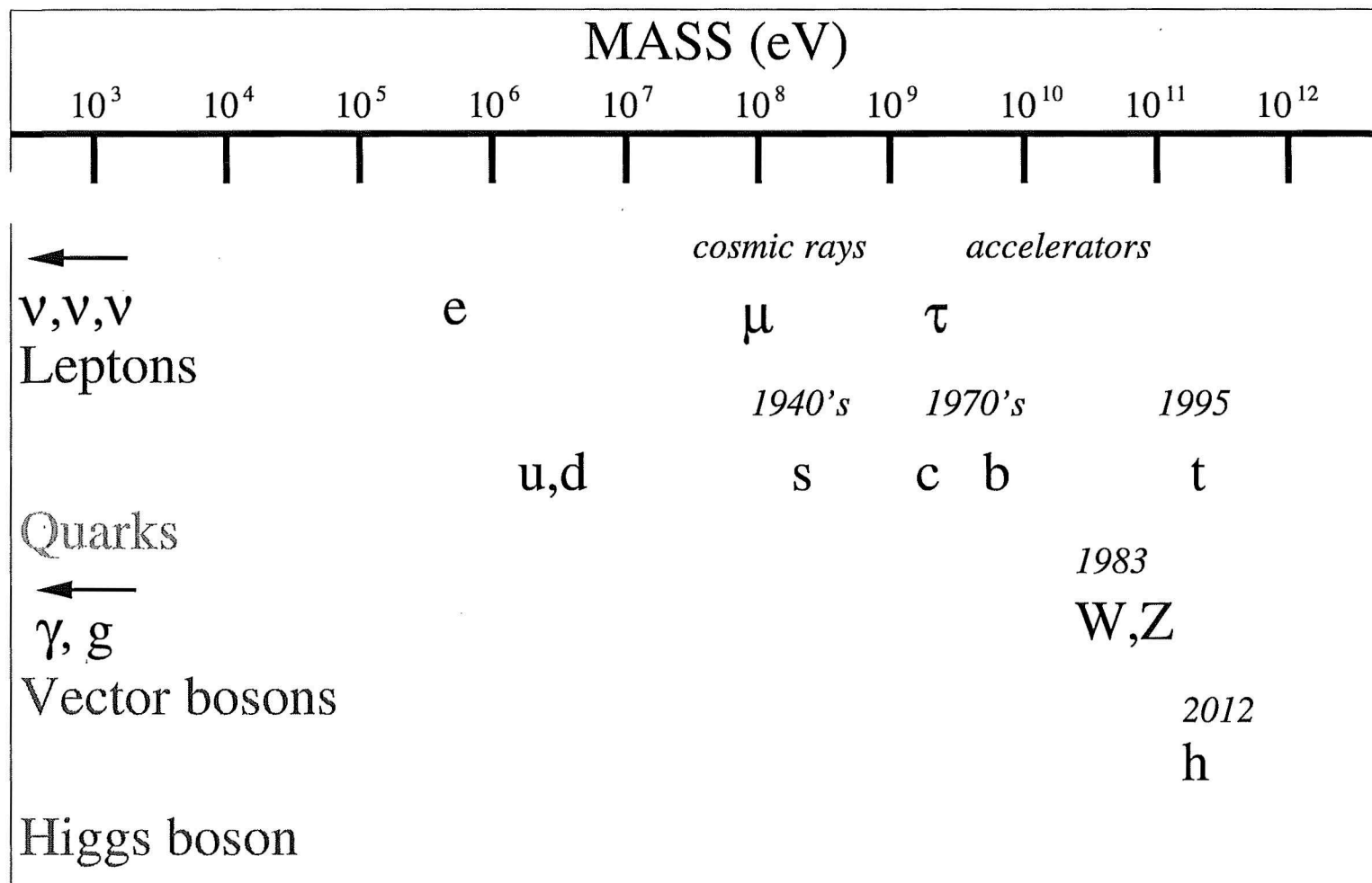
[show chart of masses]

Standard model particles

6 quarks
6 leptons

vector bosons (γ, g, W^+, W^-, Z^0)

scalar boson (Higgs)



gravity mediated by gravitational fields
governed by Einstein's equations of GR.

[These eqns have wavelike solutions
 $\Rightarrow$ gravitational wave discovered by LIGO]

Expect quantize $\rightarrow$ gravitons (massless, tensor bosons)
 Single gravitons Not (yet) detectible.

Also quantum theory of gravity is not yet satisfactory

Fortunately, gravity is irrelevant on the scale of
 elementary particles.

open problem. quantum gravity
 (string theory)

BCM? experimental
 • measure neutrinos
 • dark matter

theoretical
 • more forces (GUT)
 • Super Symmetry